

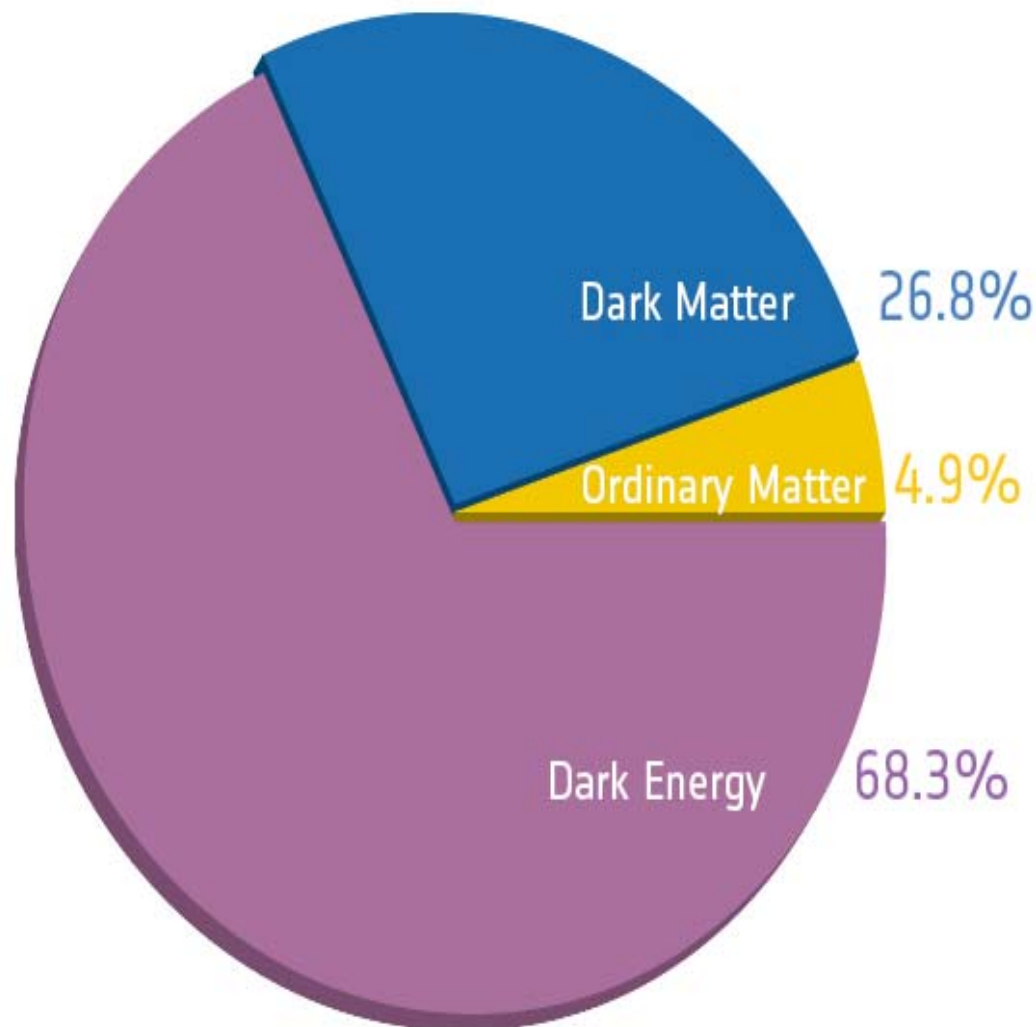
Nuclear and Particle Physics

An introduction

Τάσος Λιόλιος
Οκτώβριος 2013

Η σύσταση του σύμπαντος

(σύμφωνα με τα δεδομένα του Planck mission-2013)



A standard periodic table of elements, color-coded by blocks: s-block (yellow), p-block (orange), d-block (green), and f-block (blue). The table includes element symbols, atomic numbers, and names. It also features labels for 'Non-Metals', 'Transition Metals', 'Rare Earths', 'Lanthanide Series', and 'Actinide Series'. The table is organized into groups (I to VIII) and periods (1 to 7).

Τα δομικά στοιχεία της συνήθους ύλης
(σε ατομικό επίπεδο - χημικά στοιχεία)

1 New Designation
IA Original Designation

1 **H**
1.0094

2 **He**
4.00260

3 **Li**
6.941

4 **Be**
9.0122

5 **B**
10.81

6 **C**
12.011

7 **N**
14.007

8 **O**
15.999

9 **F**
18.998

10 **Ne**
20.179

11 **Na**
22.990

12 **Mg**
24.305

13 **Al**
26.982

14 **Si**
28.086

15 **P**
30.974

16 **S**
32.06

17 **Cl**
35.453

18 **Ar**
39.948

19 **K**
39.098

20 **Ca**
40.08

21 **Sc**
44.956

22 **Ti**
47.88

23 **V**
50.942

24 **Cr**
51.996

25 **Mn**
54.938

26 **Fe**
55.847

27 **Co**
58.933

28 **Ni**
58.69

29 **Cu**
63.546

30 **Zn**
65.39

31 **Ga**
69.72

32 **Ge**
72.59

33 **As**
74.922

34 **Se**
78.96

35 **Br**
79.904

36 **Kr**
83.80

37 **Rb**
85.468

38 **Sr**
87.62

39 **Y**
88.906

40 **Zr**
91.224

41 **Nb**
92.906

42 **Mo**
95.94

43 **Tc**
(98)

44 **Ru**
101.07

45 **Rh**
102.91

46 **Pd**
106.42

47 **Ag**
107.87

48 **Cd**
112.41

49 **In**
114.82

50 **Sn**
118.71

51 **Sb**
121.75

52 **Te**
127.60

53 **I**
126.91

54 **Xe**
131.29

55 **Cs**
132.91

56 **Ba**
137.33

57 **La**
138.91

58 **Ce**
140.12

59 **Pr**
140.91

60 **Nd**
144.24

61 **Pm**
(145)

62 **Sm**
150.36

63 **Eu**
151.96

64 **Gd**
157.25

65 **Tb**
158.93

66 **Dy**
162.50

67 **Ho**
164.93

68 **Er**
167.26

69 **Tm**
168.93

70 **Yb**
173.04

71 **Lu**
174.97

72 **Hf**
178.49

73 **Ta**
180.95

74 **W**
183.85

75 **Re**
186.21

76 **Os**
190.2

77 **Ir**
192.22

78 **Pt**
195.08

79 **Au**
196.97

80 **Hg**
200.59

81 **Tl**
204.38

82 **Pb**
207.2

83 **Bi**
208.98

84 **Po**
(209)

85 **At**
(210)

86 **Rn**
(222)

87 **Fr**
(223)

88 **Ra**
226.03

89 **Ac**
227.03

90 **Th**
232.04

91 **Pa**
231.04

92 **U**
238.03

93 **Np**
237.05

94 **Pu**
(244)

95 **Am**
(243)

96 **Cm**
(247)

97 **Bk**
(247)

98 **Cf**
(251)

99 **Es**
(253)

100 **Fm**
(257)

101 **Md**
(258)

102 **No**
(259)

103 **Lr**
(260)

104 **Unq**
(261)

105 **Unp**
(262)

106 **Unh**
(263)

107 **Uns**
(262)

108 **Uno**
(265)

109 **Une**
(266)

110 **Uun**
(267)

111 **Ubu**
(267)

112 **Ubn**
(267)

113 **Ufl**
(267)

114 **Ulu**
(267)

115 **Uts**
(267)

116 **Utt**
(267)

117 **Utt**
(267)

118 **Utt**
(267)

119 **Utt**
(267)

120 **Utt**
(267)

121 **Utt**
(267)

122 **Utt**
(267)

123 **Utt**
(267)

124 **Utt**
(267)

125 **Utt**
(267)

126 **Utt**
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127 **Utt**
(267)

128 **Utt**
(267)

129 **Utt**
(267)

130 **Utt**
(267)

131 **Utt**
(267)

132 **Utt**
(267)

133 **Utt**
(267)

134 **Utt**
(267)

135 **Utt**
(267)

136 **Utt**
(267)

137 **Utt**
(267)

138 **Utt**
(267)

139 **Utt**
(267)

140 **Utt**
(267)

141 **Utt**
(267)

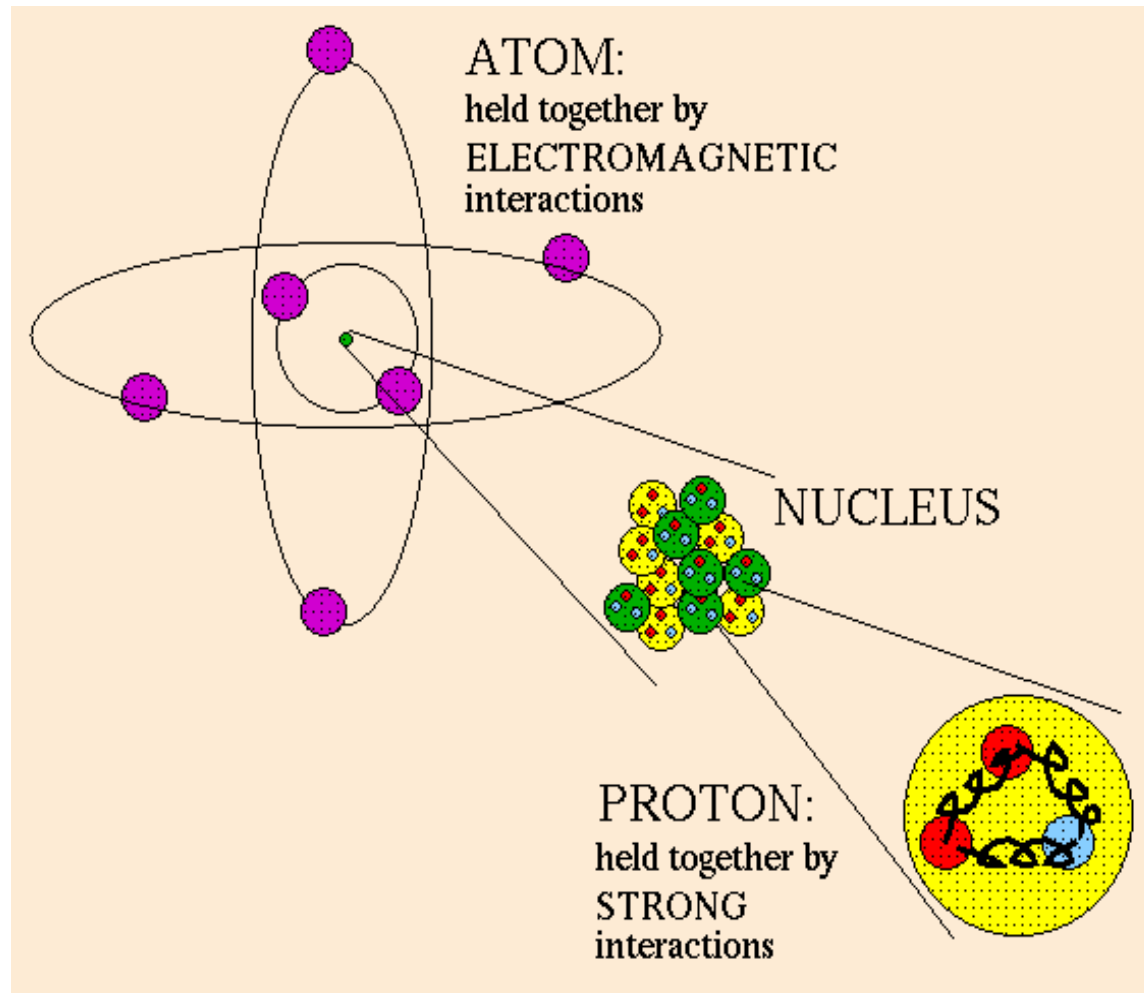
142 **Utt**
(267)

143 **Utt**
(267)

144 **Utt**
(267)

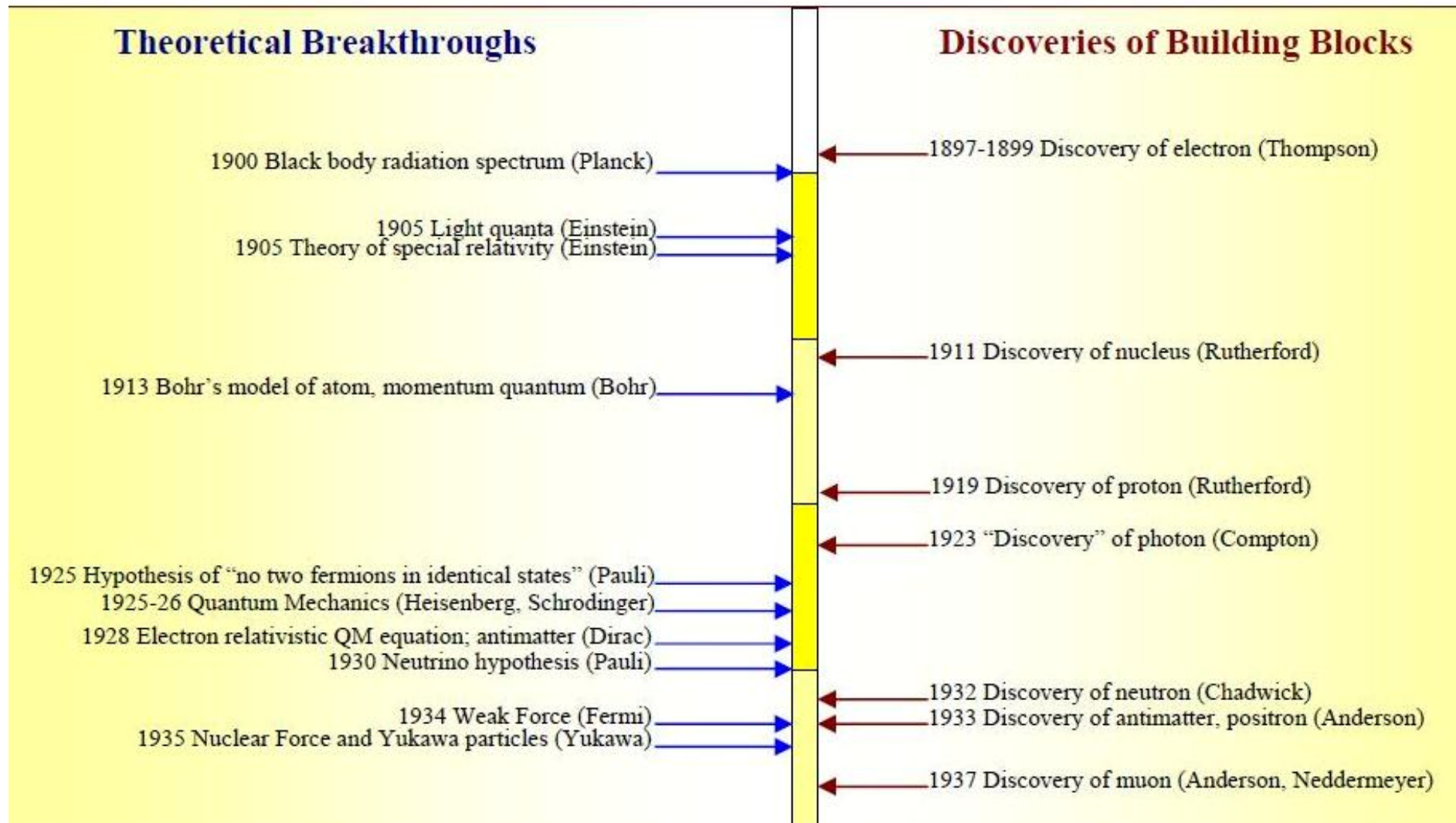
145 **Utt**
(267)

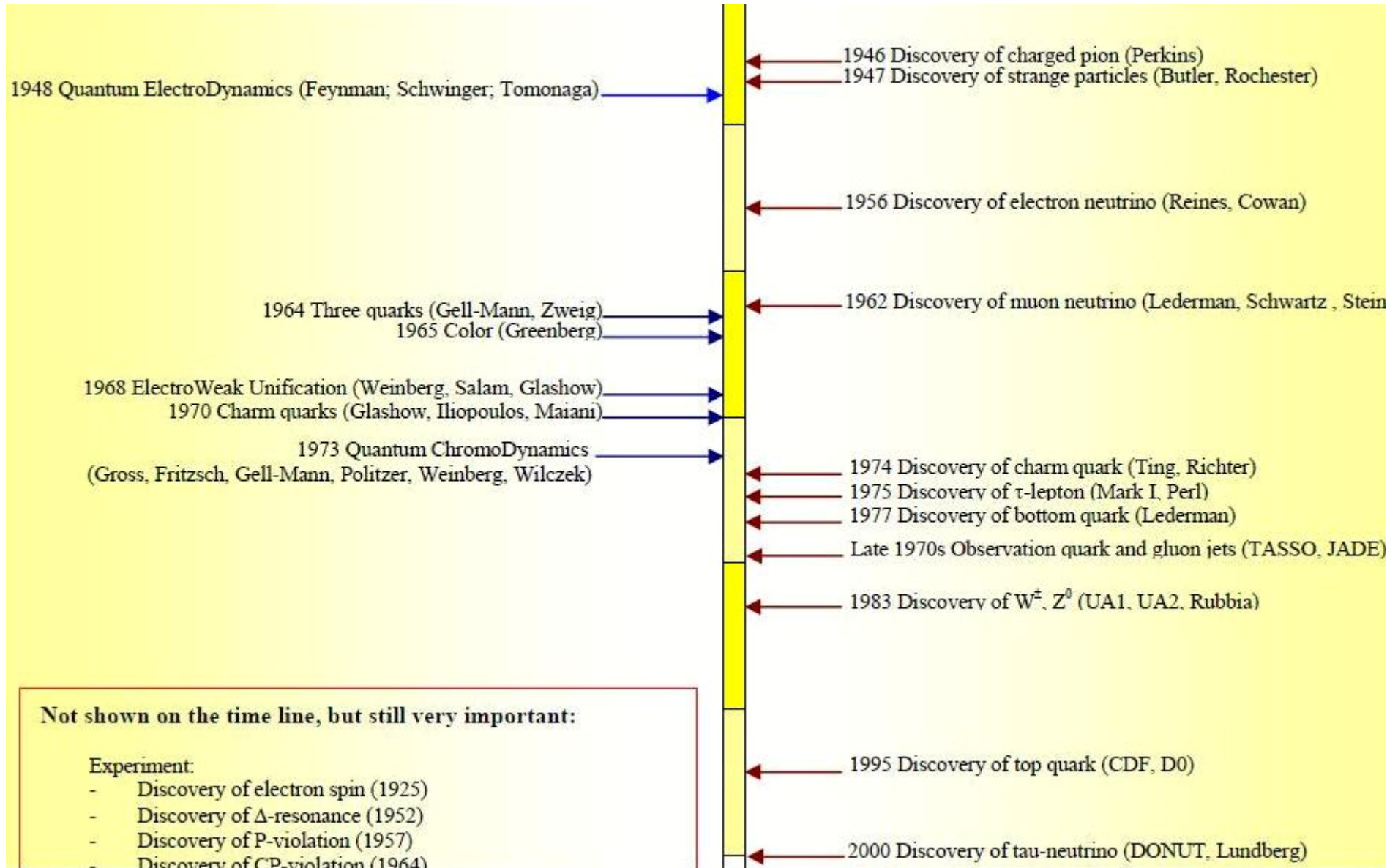
Ατομική και υπο-ατομική δόμηση της ύλης



Πώς τα ξέρουμε αυτά; Ακολουθεί λίγη ιστορία...

20th Century Elementary Particle Physics Timeline

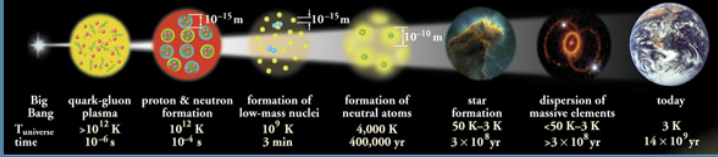




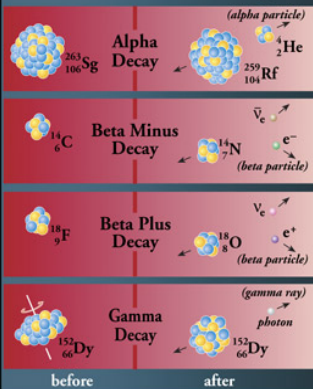
Nuclear Science

Expansion of the Universe

After the Big Bang, the universe expanded and cooled. At about 10^{-4} second, the universe consisted of a soup of quarks, gluons, electrons, and neutrinos. When the temperature of the Universe, $T_{universe}$, cooled to about 10^9 K, this soup coalesced into protons, neutrons, and electrons. As time progressed, some of the protons and neutrons formed deuterium, helium, and lithium nuclei. Still later, electrons combined with protons and these low-mass nuclei to form neutral atoms. Due to gravity, clouds of atoms contracted into stars, where hydrogen and helium fused into more massive chemical elements. Exploding stars (supernovae) form the most massive elements and disperse them into space. Our earth was formed from supernova debris.



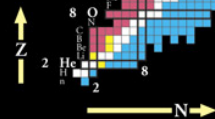
Radioactivity



Radioactive decay transforms a nucleus by emitting different particles. In **alpha** decay, the nucleus releases a ^4_2He nucleus—an alpha particle. In **beta** decay, the nucleus either emits an electron and antineutrino (or a positron and neutrino) or captures an atomic electron and emits a neutrino. A positron is the antiparticle of the electron. Antimatter is composed of anti-particles. Both alpha and beta decays change the original nucleus into a nucleus of a different chemical element. In **gamma** decay, the nucleus lowers its internal energy by emitting a photon—a gamma ray. This decay does not modify the chemical properties of the atom.

Chart of the Nuclides

The Chart of the Nuclides presents in graphic form all known nuclei with atomic number, Z , and neutron number, N . Each nuclide is represented by a box colored according to its predominant decay mode. Magic numbers (N or $Z = 2, 8, 20, 28, 50, 82$ and 126) are indicated by a rectangle on the chart. They correspond to major closed shells and show regions of greater nuclear binding energy.



Color Key

- Stable
- Spontaneous fission
- Alpha particle emission
- Beta minus emission
- Beta plus emission or electron capture

www.CPEPweb.org

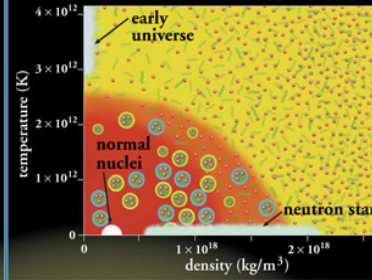
Nuclear Science is the study of the structure, properties, and interactions of the atomic nuclei. Nuclear scientists calculate and measure the masses, shapes, sizes, and decays of nuclei at rest and in collisions. They ask questions, such as: Why do nucleons stay in the nucleus? What combinations of protons and neutrons are possible? What happens when nuclei are compressed or rapidly rotated? What is the origin of the nuclei found on Earth?

Legend

- electron (e^-)
- quark
- gluon field
- proton
- positron (e^+)
- gluon
- neutrino (ν)
- antineutrino ($\bar{\nu}$)
- photon (γ)

$A_{\text{mass number}} = 14$
 $Z_{\text{atomic number}} = 6$
 $N_{\text{neutron number}} = A - Z$

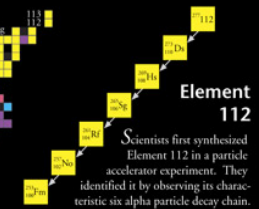
Phases of Nuclear Matter



Nuclear matter can exist in several phases. When collisions excite nuclei, individual protons and neutrons may evaporate from the nuclear fluid. At sufficiently high temperature or density, a gas of nucleons (red background) forms. At even more extreme conditions, individual nucleons may cease to have meaningful identities, merging into the quark-gluon plasma (yellow background). Current data provide hints that physicists have glimpsed the quark-gluon plasma.

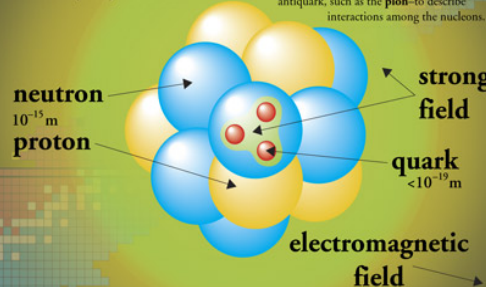
Unstable Nuclei

Stable nuclides form a narrow white band on the Chart of the Nuclides. Scientists produce unstable nuclides far from this band and study their decays, thereby learning about the extremes of nuclear conditions. In its present form, this chart contains about 2500 different nuclides. Nuclear theory predicts that there are at least 4000 more to be discovered with $Z \leq 113$.

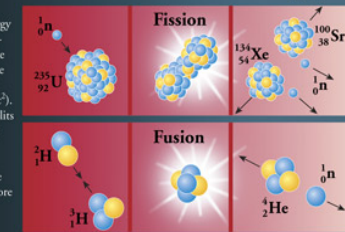


The Nucleus

$(1-10) \times 10^{-15}$ m



Nuclear Energy



In the early stages of stellar evolution of our sun and other stars, hydrogen fuses to form helium, releasing energy in the form of photons (light) and neutrinos. During the later stages of stellar evolution, more massive nuclei up to and beyond uranium are synthesized by fusion. By measuring the number of neutrinos that come from the Sun, scientists recently have demonstrated that neutrinos must have a mass greater than zero.

Applications



Radioactive Dating

Naturally occurring radioactive isotopes such as ^{14}C are used to date objects that were once living, such as wood. For example, from a study of artifacts found at the site, scientists determined that Stonehenge was built nearly 4,000 years ago.



Smoke Detectors

Many smoke detectors use a small amount of the alpha emitter ^{241}Am to ionize the air. Smoke entering the detector reduces the current and sets off the alarm.



Nuclear Medicine

Radioactive isotopes, such as $^{99\text{m}}\text{Tc}$, ^{67}Ga , and ^{18}F , are commonly used in the diagnosis and treatment of disease. Positron emitters such as ^{18}F are used in Positron Emission Tomography (PET) to generate images of brain activity.



Magnetic Resonance Imaging

Magnetic Resonance Imaging (MRI) makes use of atomic transitions involving the magnetic field of a nucleus to study the local chemical environment. This technique accurately maps the density of hydrogen to produce three-dimensional images of the human body.

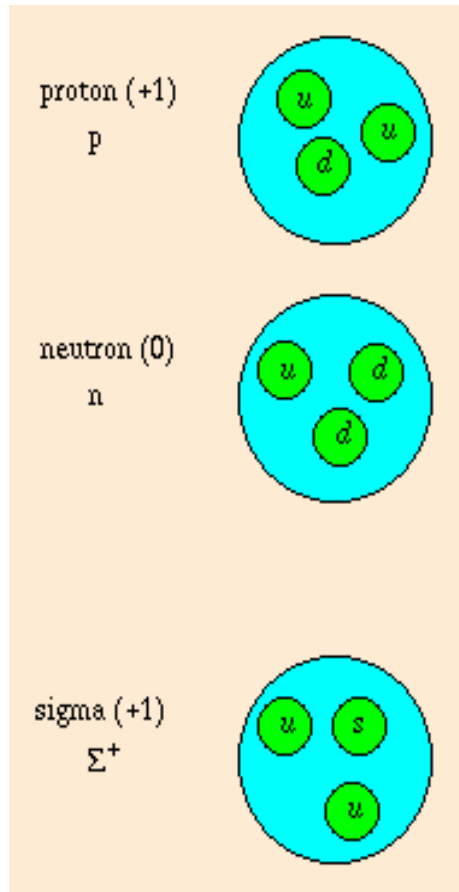


Nuclear Reactors

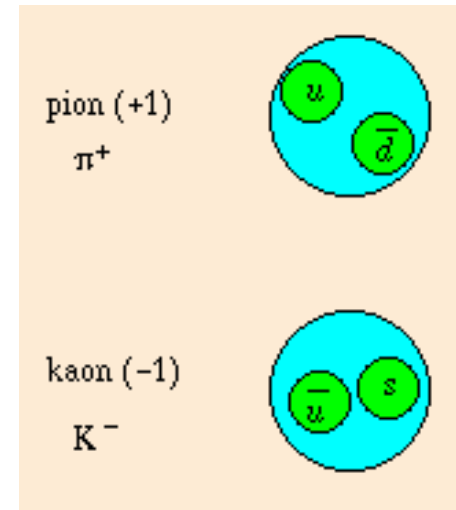
Nuclear reactors use the fusion of ^{235}U or ^{239}Pu nuclei to produce electric power. Reactors and most other nuclear applications generate radioactive waste; disposal of this waste is a subject of current research.

Astrophysical pictures courtesy NASA/JPL/Caltech and AURA/STScI.

Στοιχειώδη σωματίδια με εσωτερική δομή – αδρόνια (strong interactions)

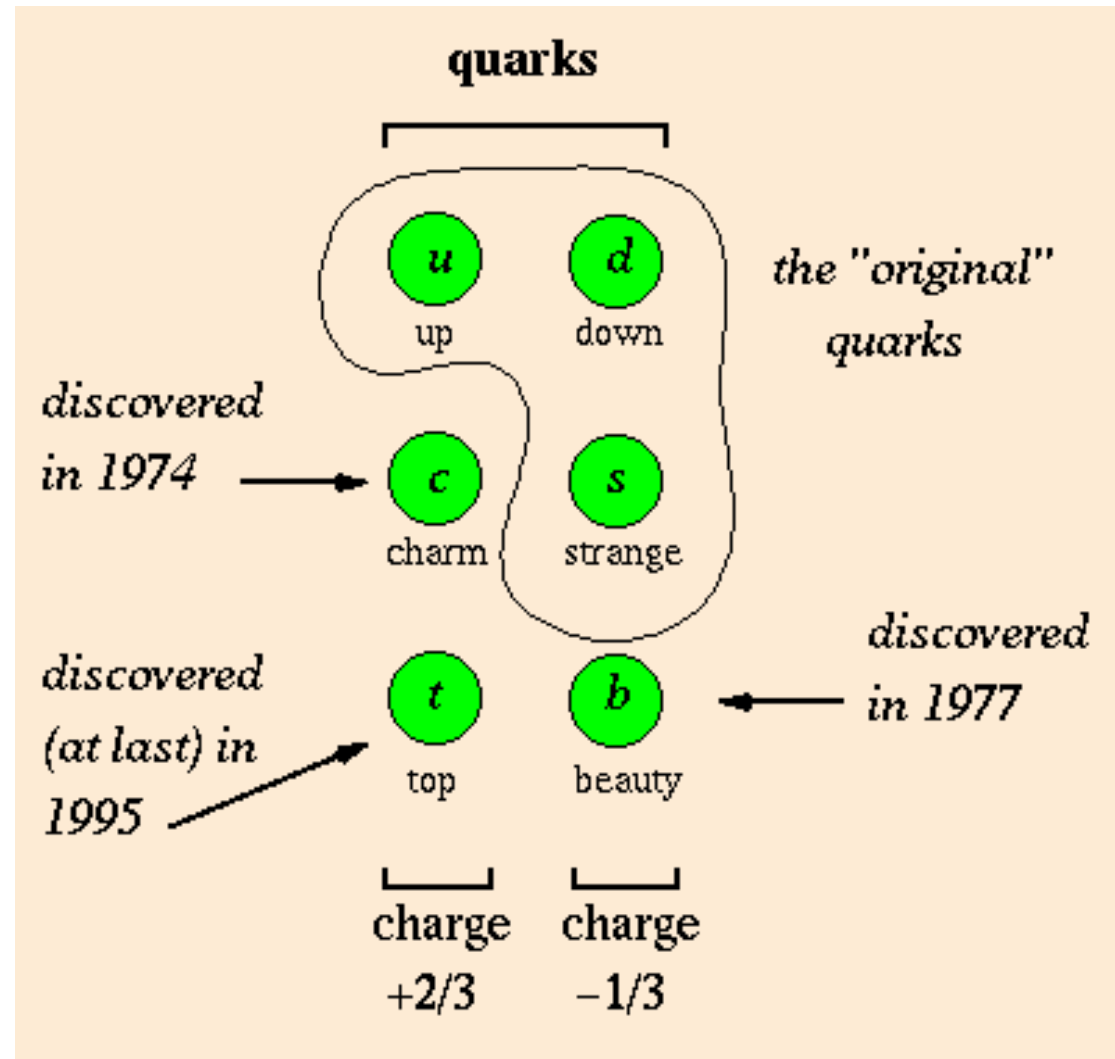


Βαρυόνια:
αδρόνια με 3 quarks ή 3 antiquark

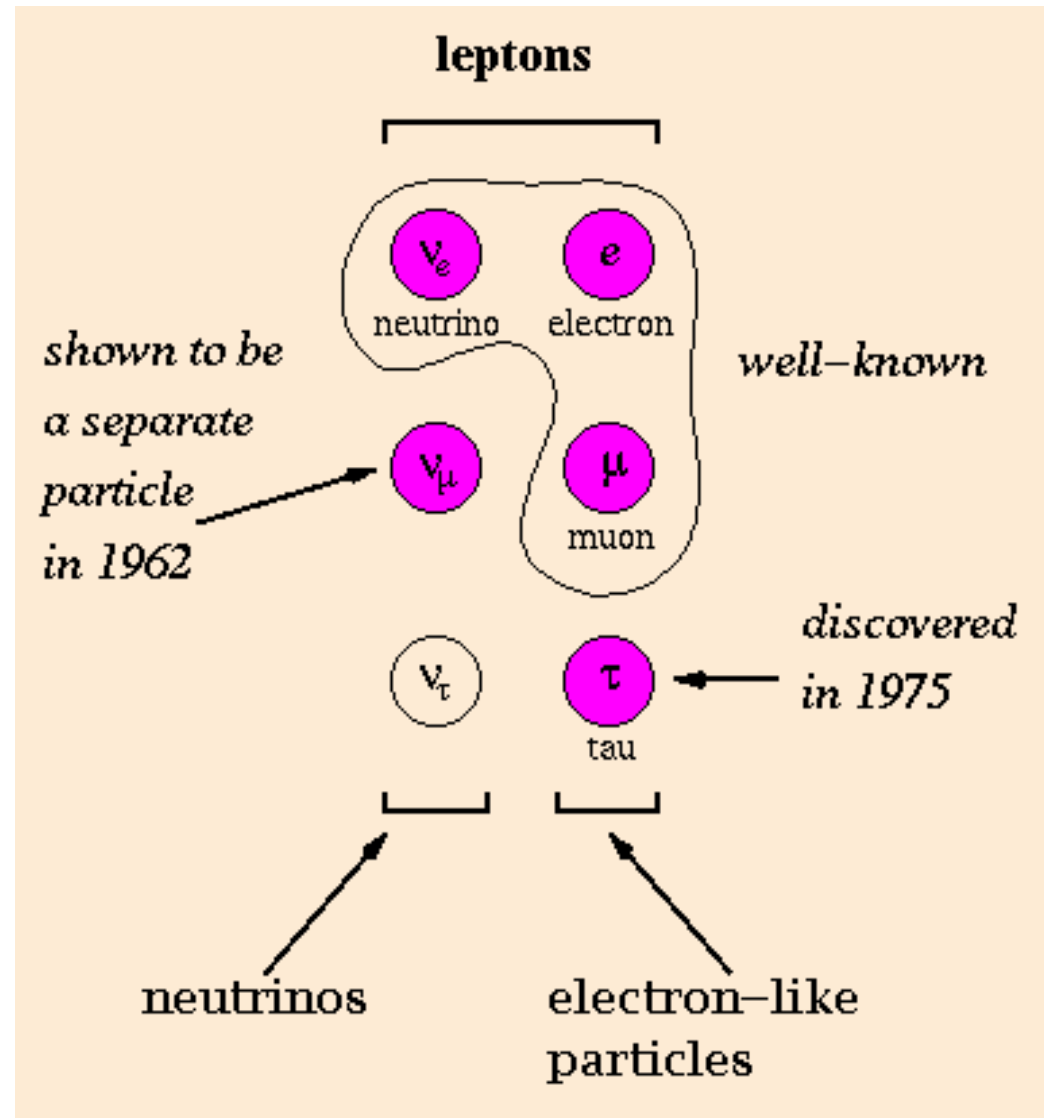


Μεσόνια:
αδρόνια με 1 quark και 1 antiquark

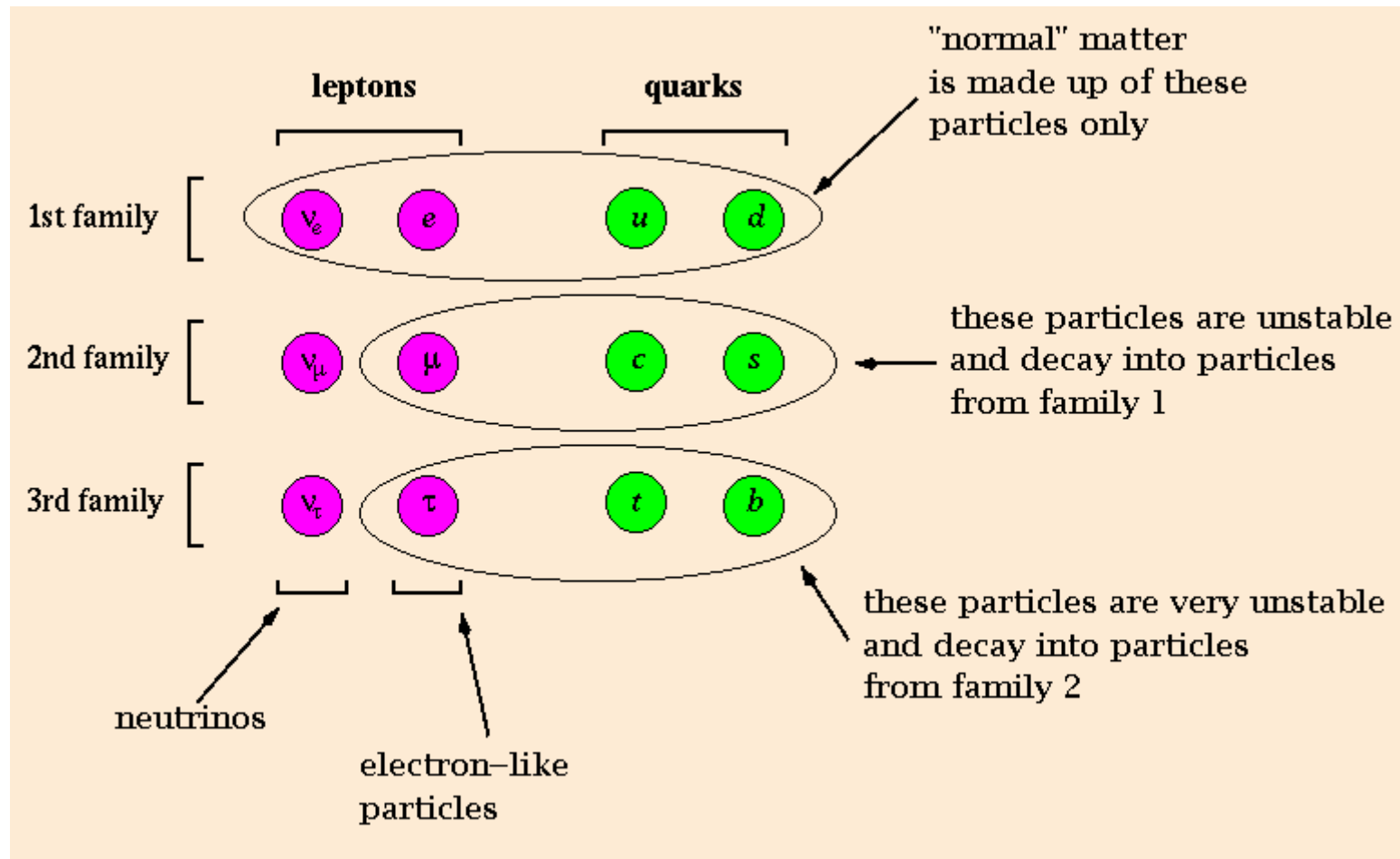
The 6 quarks



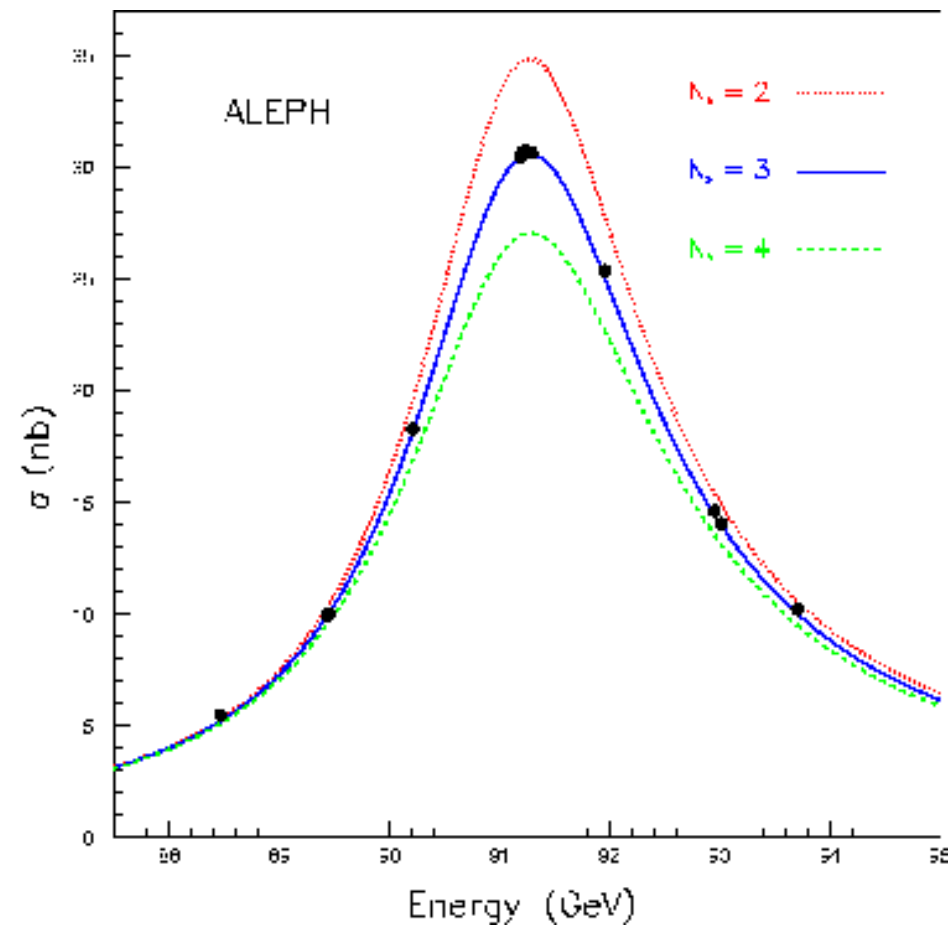
The 6 leptons



λεπτόνια και quark σε 3 οικογένειες



Measurement of Z hadronic decays, versus energy,
with predicted rates if there were 1, 2 or 3 types of neutrino
(ALEPH collaboration)



Τα δομικά στοιχεία της ύλης

(σε υποατομικό - υποπυρηνικό επίπεδο)

